

DEPARTMENTAL SEMINAR **Department of Astrophysics and High Energy Physics**

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ONLINE/ FERMION

SPEAKER

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TITLE OF THE TALK

Magnetic Fields in Massive Star-forming Regions (MagMaR)

ABSTRACT

The importance of magnetic fields (B-fields) in the high-mass star-forming process is a long-standing question. Indeed, several observational properties in high-mass star-forming regions (e.g., high/low level of core fragmentation) are frequently explained invoking B-fields, despite the lack of direct evidence of their presence or their importance with respect to turbulence and gravity. Some efforts have been attempted to address the importance of B-fields by observing statistically significant samples, for example, with SMA and CARMA, but not so far with ALMA. Here, we introduce the first ALMA survey, Magnetic Fields in Massive Star-forming Regions (MagMaR). In MagMaR, 30 high-mass star-forming regions have been observed at 1.2 mm, resulting in -0.3" resolution (-1000 au). A large variety of B-field morphologies is detected: (1) "simple" spiral- or hourglass-like, with little/no fragmentation; (2) filamentary, with B-field vectors sometimes parallel to the elongated dust emission and characterized with aligned fragmentation; (3) complex B-field morphology with highly clustered fragmentation. The initial findings of the survey and the results of some case studies will be presented in this talk.

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