

# Non-supersymmetric D3-brane: complexity and information metric

(Based on: arXiv:1807.06361 with Aranya Bhattacharya)

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- 1 Motivation and Introduction
- 2 Non-supersymmetric D3 brane
- 3 Complexity and Information metric
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# Motivation and Introduction

- In the last few years there has been a growing interest to find relations between quantum information theory and the geometry or gravity.
- Apparently there should not be any connection. Here one side is purely quantum mechanical without gravity and the other side is a gravitational theory.
- The bigger picture is to understand the black holes better. Particularly, the information loss paradox. Necessary to get a consistent theory of quantum gravity.
- This paradox appeared to have been resolved by the idea of stretched horizon and black hole complementarity proposed by Susskind (also by Preskill and 't Hooft).

- However, a few years ago it has been pointed out by AMPS [A. Almheiri, D. Marolf, J. Polchinski, J. Sully, JHEP02 (2013) 62] that Susskind idea does not resolve the paradox. It leads to a dramatic event (another paradox) the so-called **firewall** on the horizon which is against the principle of general relativity.



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- A possible resolution of **firewall** paradox was suggested by Daniel Harlow and Patrick Hayden [D. Harlow and P. Hayden, JHEP06 (2013) 85] . They first introduced ‘computational complexity’, a concept borrowed from information theory.
- The **firewall** argument depends on an observer’s ability to decode information from the outgoing Hawking radiation near the horizon. Using complexity Harlow and Hayden showed that the number of steps required to decode the outgoing information increases exponentially as  $e^S$  ( $S$  is the entropy of the black hole).

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- Therefore, there is a need to understand various concepts in information theory, such as complexity, information metric (information geometry) to better understand the quantum nature of our space-time.
- More formally the **quantum state complexity** tells us how difficult it is to go from a reference state  $|\psi_0\rangle$  to a final state  $|\psi\rangle$  by some series of unitary transformation.

$$|\psi\rangle = U|\psi_0\rangle$$

Complexity is the minimum number of unitary transformations (or gates) to obtain  $U$

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- Moreover, when this analysis was applied to a certain tensor network (called **Multiscale Entanglement Renormalization Ansatz or MERA**) used to obtain ground state of certain critical physical theories, it gives a tensor network structure of AdS time slice.



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- Moreover, when this analysis was applied to a certain tensor network (called **Multiscale Entanglement Renormalization Ansatz or MERA**) used to obtain ground state of certain critical physical theories, it gives a tensor network structure of AdS time slice.
- This was the primary motivation for the **Complexity = Volume (CV)** and **Complexity = Action (CA)** proposal of Susskind et. al. (for AdS-Schwarzschild black hole) [A. R. Brown, D. A. Roberts, L. Susskind and B. Swingle, *PRL*116 (2016) 191301]

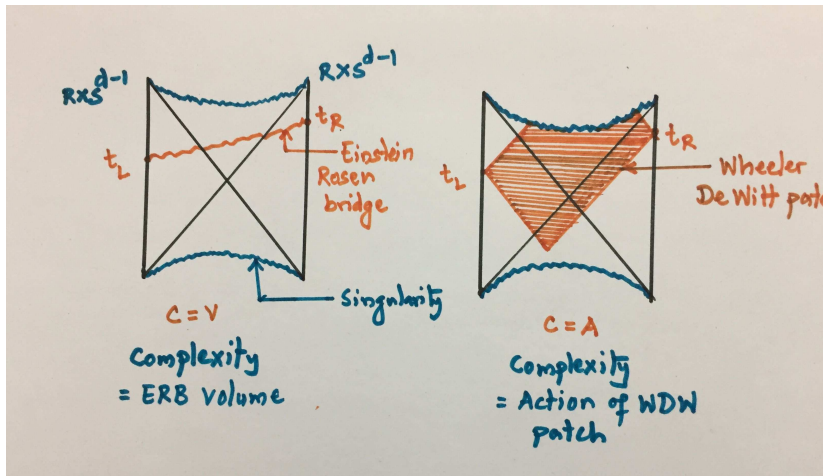
- According to Susskind,
  1. **CV conjecture**: Holographic Complexity (of a CFT state) = Maximum volume of Einstein Rosen Bridge connecting the two black holes (two sided black holes).

$$C_V(\Sigma) = \max_{\Sigma=\partial\mathcal{B}} \left[ \frac{V_{\mathcal{B}}}{G_N R} \right]$$

where  $\mathcal{B}$  is the codimension-one bulk surface,  $\partial\mathcal{B}$  is its boundary =  $\Sigma$  and  $V_{\mathcal{B}}$  is the volume.  $G_N$ , Newton's constant and  $R$  is the AdS radius.

2. **CA conjecture**: Holographic Complexity = Classical action of the certain region of the bulk called the Wheeler-De Witt Patch

$$C_A(\Sigma) = \frac{I_{WDW}}{\pi\hbar}$$



- We will be computing complexity (**Subregion Complexity**) using another definition proposed by Alishahiha [M. Alishahiha, Phys. Rev. D92 (2015) 126009] motivated by Ryu-Takayanagi [S. Ryu and T. Takayanagi, PRL 96 (2006) 181602] formula.

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- Ryu and Takayanagi proposed how to compute another information theoretic quantity called the **entanglement entropy** for a bipartite system holographically.

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- Ryu and Takayanagi proposed how to compute another information theoretic quantity called the **entanglement entropy** for a bipartite system holographically.
- For a system consisting of two parts  $A$  and  $B$  (complement of  $A$ ), EE of a subsystem  $A$  is the von Neumann entropy and is defined as

$$S_A = -\text{Tr}(\rho_A \log \rho_A)$$

where  $\rho_A = \text{Tr}_B(\rho_{\text{tot}})$ .

- According to Ryu-Takayanagi: the EE can be calculated holographically as

$$S_E(A) = \frac{\text{Area}(\gamma_A^{\min})}{4G_N}$$

$\gamma_A^{\min}$  is the minimal codimension 2 surface in  $\text{AdS}_{d+2}$  space with  $\partial\gamma_A^{\min} = \partial A$  and  $G_N$  is the  $(d+2)$ -dimensional Newton's constant.

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- Alishahiha made an observation: The Ryu-Takayanagi minimal codimension two area naturally gives rise to another unique quantity on the gravity side. It is the **regularized volume** enclosed by the area. By dimensionality it must be some form of complexity. So, he defined:

$$\text{Subregion complexity } (C_V) = \frac{V(\gamma)}{8\pi R G_N}$$

$\gamma$  is the same Ryu-Takayanagi surface as before.



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- Here we show the computation of **subregion complexity** for a particular gravitational system given by the **decoupled geometry of non-supersymmetric D3 brane**. Also we compute the **Fisher information metric**.
- To introduce the notion of Fisher information metric, we mention that if  $\rho_\lambda$  is a density matrix parametrized by a parameter  $\lambda$  then a new density matrix  $\rho_{\lambda+\delta\lambda}$  can be obtained by perturbing it by  $\delta\lambda$ . Thus,

$$\rho_{\lambda+\delta\lambda} = \rho_\lambda + (\delta\lambda)\rho_1 + \frac{1}{2}(\delta\lambda)^2\rho_2 + \dots$$

Then a notion of distance metric may be defined as,

$$G_{\lambda\lambda} = \langle \delta\rho\delta\rho \rangle_{\lambda\lambda} = \frac{1}{2}\text{Tr} \left( \delta\rho \frac{d}{d(\delta\lambda)} \log(\rho_\lambda + \delta\lambda\delta\rho) \Big|_{\delta\lambda=0} \right)$$

- This is obtained by calculating the **relative entropy** between the two adjacent density matrices as [N. Lashkari and M. van Raamsdonk, JHEP 04 (2016) 153]

$$S(\rho_{\lambda+\delta\lambda}||\rho_{\lambda})_{\lambda^2} = (\delta\lambda)^2 \langle \delta\rho \delta\rho \rangle_{\lambda\lambda} = (\delta\lambda)^2 G_{\lambda\lambda}$$

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- For the two adjacent density matrices one can define a quantity called the **Fidelity** defined as,

$$\mathcal{F} = \text{Tr} \sqrt{\sqrt{\rho_{\lambda}} \rho_{\lambda+\delta\lambda} \sqrt{\rho_{\lambda}}}$$

Again a quantum distance between the same two states can be defined as (also known as **Fidelity Susceptibility**) [S. -J. Gu, IJMPB 24 (2010) 4371]

$$\tilde{G}_{\lambda\lambda} = \partial_{\lambda}^2 \mathcal{F}$$

- For the kind of states we will be considering (CFT vacuum state dual to pure  $\text{AdS}_5$  and its marginal deformation by the dilaton),  $G_{\lambda\lambda}$  and  $\tilde{G}_{\lambda\lambda}$  can be shown to be the same.

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- The **Fidelity** can be calculated holographically essentially from the **subregion complexity** proposed by Alishahiha  $\Rightarrow$  the regularized volume of the Ryu-Takayanagi surface [S. Banerjee, J. Erdmenger and D. Sarkar, JHEP 08 (2018) 001] .

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- The **Fidelity** can be calculated holographically essentially from the **subregion complexity** proposed by Alishahiha  $\Rightarrow$  the regularized volume of the Ryu-Takayanagi surface [S. Banerjee, J. Erdmenger and D. Sarkar, JHEP 08 (2018) 001] .
- This is the difference of volumes of the excited state (perturbed state upto second order) and the ground state (pure AdS<sub>5</sub>) i.e.,

$$\mathcal{F} = C_d \left( V^{(m^2)} - V^{(0)} \right)$$

where  $m$  is the perturbation parameter and  $C_d$  is a  $d$ -dimensional constant. The Fisher information metric then can be obtained as,

$$G_{mm} = \partial_m^2 \mathcal{F}$$

# Non-supersymmetric D3 brane



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- Standard D3 brane is supersymmetric (BPS). Here we consider a non-supersymmetric version of D3 brane. This is also a solution of low energy Type IIB string theory as BPS D3 brane. [B. Zhou and C.J. Zhu, hep-th/9905146; P. Brax, G. Mandal and Y. Oz, PRD 63 (2001) 064008; J. X. Lu and S. Roy, JHEP 02 (2005) 001.]

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- The solution (with metric in Einstein frame) in some suitable coordinates take the form,

$$ds^2 = F_1(\rho)^{-\frac{1}{2}} G(\rho)^{-\frac{\delta_2}{8}} \left[ -G(\rho)^{\frac{\delta_2}{2}} dt^2 + \sum_{i=1}^3 (dx^i)^2 \right] + F_1(\rho)^{\frac{1}{2}} G(\rho)^{\frac{1}{4}} \left[ \frac{d\rho^2}{G(\rho)} + \rho^2 d\Omega_5^2 \right]$$

$$e^{2\phi} = G(\rho)^{-\frac{3\delta_2}{2} + \frac{7\delta_1}{4}}, \quad F_{[5]} = \frac{1}{\sqrt{2}} (1 + *) Q \text{Vol}(\Omega_5).$$

- The functions  $G(\rho)$  and  $F_1(\rho)$  are defined as,

$$G(\rho) = 1 + \frac{\rho_0^4}{\rho^4}, \quad F_1(\rho) = G(\rho)^{\frac{\alpha_1}{2}} \cosh^2 \theta - G(\rho)^{-\frac{\beta_1}{2}} \sinh^2 \theta$$

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- There are several parameters  $(\delta_1, \delta_2, \alpha_1, \beta_1, \theta, \rho_0, Q)$  in the solution. They satisfy certain relations. For simplicity we will choose  $\alpha_1 = \beta_1 = 1$ . The other relations are,

$$42\delta_2^2 + 49\delta_1^2 - 84\delta_1\delta_2 = 24, \quad Q = 2\rho_0^4 \sinh 2\theta$$

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- So, there are three independent parameters  $\delta_2, \rho_0$  and  $\theta$ .  $Q$  is the charge of D3 brane.  $F_1$  now reduces to

$$F_1(\rho) = \left( 1 + \frac{\rho_0^4 \cosh^2 \theta}{\rho^4} \right) G(\rho)^{-\frac{1}{2}}$$

In the double scaling limit  $\rho_0 \rightarrow 0, \theta \rightarrow \infty$  such that  $\rho_0^4 \cosh^2 \theta = R_1^4 = \text{fixed}$ , we recover the standard BPS D3 brane from the non-supersymmetric D3 brane solution given before. In this case  $G(\rho) \rightarrow 1$ .

- The solution

$$ds^2 = F_1(\rho)^{-\frac{1}{2}} G(\rho)^{-\frac{\delta_2}{8}} \left[ -G(\rho)^{\frac{\delta_2}{2}} dt^2 + \sum_{i=1}^3 (dx^i)^2 \right] + F_1(\rho)^{\frac{1}{2}} G(\rho)^{\frac{1}{4}} \left[ \frac{d\rho^2}{G(\rho)} + \rho^2 d\Omega_5^2 \right]$$

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- Also  $\delta_2 = 0$  corresponds to zero temperature solution. That is also non-supersymmetric (unlike BPS D3 brane).

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- Decoupling limit is zooming into the region [K. Nayek and S. Roy, PLB 766 (2017) 192] :

$$\rho \sim \rho_0 \ll \rho_0 \cosh^{\frac{1}{2}} \theta \quad \Rightarrow \quad \theta \rightarrow \infty$$



- The solution reduces to:

$$ds^2 = \frac{\rho^2}{R_1^2} G(\rho)^{\frac{1}{4} - \frac{\delta_2}{8}} \left[ -G(\rho)^{\frac{\delta_2}{2}} dt^2 + \sum_{i=1}^3 (dx^i)^2 \right] + \frac{R_1^2}{\rho^2} \frac{d\rho^2}{G(\rho)} + R_1^2 d\Omega_5^2$$

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- The above solution is asymptotically AdS<sub>5</sub> since for  $\rho \rightarrow \infty$ ,  
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- For non-zero  $\delta_2$ , the background has a temperature [Y. Kim, B. H. Lee and S. J. Sin, JHEP09 (2007) 105]

$$T_{\text{nonsusy}} = \left( \frac{-\delta_2}{2} \right)^{\frac{1}{4}} \frac{1}{\pi \rho_0 \cosh \theta}$$

Here  $-2 \leq \delta_2 \leq 0$  and for  $\delta_2 = -2$ , it reduces to standard AdS<sub>5</sub> black hole temperature.



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- We make two coordinate changes as,

$$\tilde{\rho}^4 = \rho^4 + \rho_0^4, \quad \text{and} \quad z = R_1^2 / \tilde{\rho}$$

# Complexity and Information metric

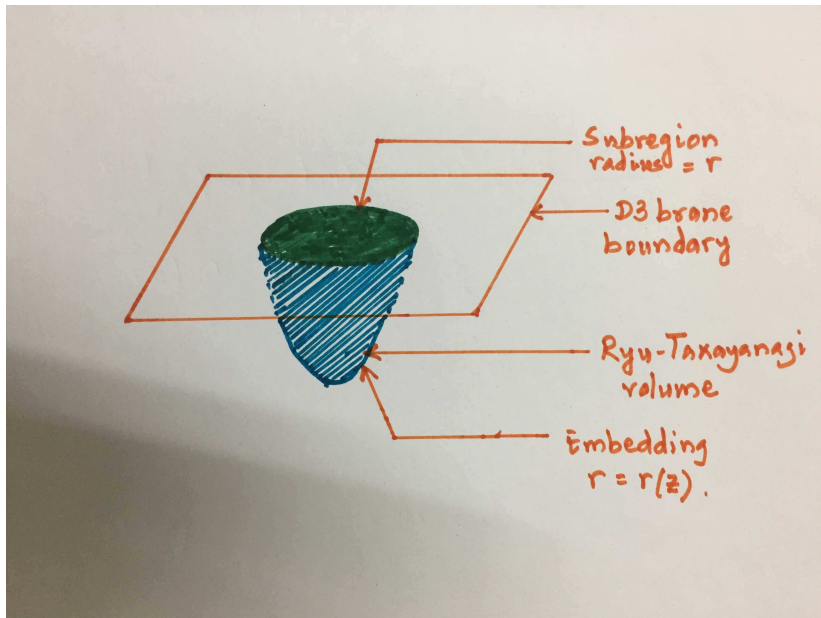
- As we mentioned before, to compute the complexity we need to calculate the Ryu-Takayanagi volume. For this we first rewrite the background in a slightly different way using some coordinate transformations.
- We make two coordinate changes as,

$$\tilde{\rho}^4 = \rho^4 + \rho_0^4, \quad \text{and} \quad z = R_1^2/\tilde{\rho}$$

- The background in  $z$  coordinate takes the form,

$$ds^2 = \frac{R_1^2}{z^2} \left[ -(1 - mz^4)^{\frac{1}{4} - \frac{3\delta_2}{8}} dt^2 + (1 - mz^4)^{\frac{1}{4} + \frac{\delta_2}{8}} \left( dr^2 + r^2 d\Omega_2^2 \right) + \frac{dz^2}{(1 - mz^4)} \right]$$

where  $m = \rho_0^4/R_1^8$  and we have written  $\sum_{i=1}^3 (dx^i)^2 = dr^2 + r^2 d\Omega_2^2$ . Note that for  $\delta_2 = -2$ , it reduces to the AdS<sub>5</sub> black hole solution.



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- To calculate  $r(z)$ , we first calculate the area of the embedded surface (shown in Fig.) using the decoupled non-susy D3 brane metric

$$\begin{aligned}
 A_{nsD3} &= \int \sqrt{g} dz d\theta d\phi \\
 &= 4\pi R_1^3 \int_{z=\epsilon}^L dz \frac{r(z)^2 (1 - mz^4)^{\frac{\delta_2}{8} - \frac{1}{4}}}{z^3} \left[ 1 + (1 - mz^4)^{\frac{5}{4} + \frac{\delta_2}{8}} r'(z)^2 \right]^{\frac{1}{2}}
 \end{aligned}$$

Since  $z = 0$  is the boundary  $\epsilon$  is an UV cut-off and  $L$  is related to the radius of the spherical subsystem  $\ell$ . They are same when  $m = 0$ .

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- $m = 0$  corresponds to pure AdS<sub>5</sub> case and  $\delta_2 = -2$  corresponds to AdS<sub>5</sub> black hole.
- We use the E-L equation of motion for  $r$  and solve it with suitable boundary condition.



- For  $m = 0$ , the solution to the equation of motion is  $r(z) = \sqrt{L^2 - z^2} = \sqrt{\ell^2 - z^2}$ , where  $r(0) = L = \ell$  has been used.

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- By replacing  $m$  by  $\eta/L^4$  and making an ansatz  $r(z) = \sqrt{L^2 - z^2} + \eta R(z)$ , we solve the equation for  $R(z)$  with BC (i)  $R(0) = L/5$  and (ii)  $\lim_{z \rightarrow L} R(z) = 0$ . This way the embedding  $r(z)$  can be obtained as

$$r(z) = \sqrt{L^2 - z^2} \left[ 1 + \eta \frac{4\delta_2 L^2 z^4 - (\delta_2 + 10)z^6 + 16L^6}{80L^4(L^2 - z^2)} \right]$$

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- Once we have the embedding we can calculate the Ryu-Takayangi volume as,

$$V_{nsD3} = \frac{4\pi R_1^4}{3} \int_{\epsilon}^L \frac{dz}{z^4} r(z)^3 \left( 1 - \eta \frac{z^4}{L^4} \right)^{\frac{3\delta_2}{16} - \frac{1}{8}}$$

- Now using the form of  $r(z)$  in the volume integral we can evaluate it upto second order in  $\eta$  as,

$$V_{0(nsD3)} = \frac{4\pi R_1^4}{3} \int_{\epsilon}^L dz \left[ \frac{(L^2 - z^2)^{3/2}}{z^4} \right] = V_{AdS}$$

$$V_{1(nsD3)} = \eta \left( \frac{\pi R_1^4}{60L^4} \right) \int_{\epsilon}^L \frac{dz}{z^4} \left[ \sqrt{(L^2 - z^2)} \left( (10 - 3\delta_2)L^2 z^4 + 4(3\delta_2 - 10)z^6 + 48L^6 \right) \right]$$

$$V_{2(nsD3)} = \eta^2 \left( \frac{\pi R_1^4}{3200L^8} \right) \int_{\epsilon}^L \frac{dz}{z^4 \sqrt{L^2 - z^2}} \left[ 32(10 - 7\delta_2)L^8 z^4 + 32(13\delta_2 - 30)L^6 z^6 \right. \\ \left. + (300 - \delta_2(13\delta_2 + 420)) L^4 z^8 - 16((\delta_2 - 65)\delta_2 + 50) L^2 z^{10} \right. \\ \left. + (\delta_2(47\delta_2 - 740) + 700) z^{12} + 512L^{12} \right]$$

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- Evaluating the integral and sending  $\epsilon \rightarrow 0$ , we get the subregion complexity using the Alishahiha proposal as,

$$C_V^{(1)} = \frac{V_{1(nsD3)} - V_{0(nsD3)}}{8\pi R_1 G_5} = 0$$

$$C_V^{(2)} = \frac{V_{2(nsD3)} - V_{0(nsD3)}}{8\pi R_1 G_5} = \left( \frac{4\pi R_1^3}{24\pi G_5} \right) \left[ \frac{3\pi (381\delta_2^2 - 6508\delta_2 - 13580)}{3276800} \right] (m\ell^4)^2 \\ = \frac{\pi R_1^3 (381\delta_2^2 - 6508\delta_2 - 13580)}{6553600 G_5} (m\ell^4)^2$$

- The holographic fidelity between two states has been given before as the difference in the volume upto a  $d$ -dimensional constant as  $\mathcal{F} = C_d \left( V^{(m^2)} - V^{(0)} \right)$ . In our case we can identify the constant to be

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$$\mathcal{F}_{nsD3} = C_4 (V_{2(nsD3)} - V_{0(nsD3)}) = \frac{32\pi^2}{4725G_5} R_1^3 m^2 \ell^8$$



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- To our surprise, this is independent of  $\delta_2$ . Therefore, it has same value even for AdS<sub>5</sub> black hole. This is independent of underlying supersymmetry of the theory.

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- The dual field theory corresponding to decoupled geometry of non-supersymmetric D3 brane is a non-conformal, non-supersymmetric QFT. This can be thought of as a perturbation of  $D = 4$ ,  $\mathcal{N} = 4$  CFT. It will be very interesting to understand how to compute complexity or information metric on the field theory side.

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- Various connections of many other quantities of quantum information theory and space-time geometry are still being revealed and hopefully, they will clarify the mysteries of quantum gravity.

# Thank You!!